

# A conservative adaptive wavelet method for the shallow-water equations on the sphere

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We introduce an innovative wavelet-based approach to adjust local grid resolution dynamically to maintain a uniform specified error tolerance. Extending the work of Dubos and Kevlahan, a wavelet multiscale approximation is used to make the Thuburn-Ringler-Skamarock-Klemp (TRiSK) model dynamically adaptive for the rotating shallow-water equations on the sphere. This article focuses on the challenges encountered when extending the adaptive wavelet method to the sphere and ensuring an efficient parallel implementation using message passing interface (MPI). The wavelet method is implemented in Fortran 95 with an emphasis on computational efficiency and scales well up to  $O(10^2)$  processors for load-unbalanced scenarios and up to at least  $O(10^3)$  processors for load-balanced scenarios. The method is verified using standard smooth test cases and a nonlinear test case proposed by Galewsky et al. The dynamical grid adaption provides compression ratios of up to 50 times in a challenging homogenous turbulence test case. The adaptive code is about three times slower per active grid point than the equivalent non-adaptive TRiSK code and about four times slower per active grid point than an equivalent spectral code. This computationally efficient adaptive dynamical core could serve as the foundation on which to build a complete climate or weather model.

Key Words: shallow-water equations; wavelets; adaptive numerical simulation; dynamical core

Received 27 March 2014; Revised 4 August 2014; Accepted 24 September 2014; Published online in Wiley Online Library

# 1. Introduction

# 1.1. Adaptive spherical shallow-water models

Geophysical flows are characterized by a wide range of time and space scales. Eddies, jets, currents and wave packets are typical features that appear locally and involve small scales. It is also necessary for global circulation models to resolve large-scale features with length-scales of thousands of kilometres. Because the location of the smallest dynamically active scales changes incessantly, an optimally efficient computational model should have a dynamically adaptive grid that tracks small-scale features and ensures that numerical errors remain below a target value. In other words, the model should automatically adapt its resolution locally where required, in order to resolve emerging small-scale features or coarsen as these features dissipate. Fixed nested and stretched grids (Krinner et al., 1997) have been used in weather forecasting and regional climate modelling. However the nonuniform grid resolution of these statically adaptive models is based on a priori knowledge of the solution, which is not possible for strongly nonlinear and non-stationary flows.

Dynamical adaptivity, where the grid is adapted automatically based on the solution and changes in time, is still a research topic in geophysical science and has not yet been incorporated

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into operational general circulation models. The book by Behrens (2009) gives an introduction and overview of adaptive modelling in atmospheric science. Pioneered by Skamarock *et al.* (1989) for weather models, dynamical adaptivity has been introduced for rotating shallow-water models by Jablonowski *et al.* (2009) (block-structured, finite-volume, latitude–longitude). Several models (interpolation-based, spectral element, cubed sphere) were compared by St-Cyr *et al.* (2008). Solutions of the shallow-water equations on statically (Ringler *et al.*, 2011) or dynamically (Bauer *et al.*, 2013) stretched unstructured meshes have also been examined. More recently, wavelets have been used for adaptive ocean modelling by Reckinger (2011), although this was a collocation method on the plane that does not conserve mass. The potential of dynamically adaptive numerical methods for global ocean and atmosphere modelling is still being explored.

# 1.2. Contributions of this work and outline

A conservative adaptive wavelet method for shallow-water equations on a regular staggered hexagonal C-grid was recently introduced by Dubos and Kevlahan (2013). This prototype method was implemented for regular planar geometry in Matlab and this demonstrated the potential of this dynamically adaptive method for simulating multiscale geophysical flows. The present work is a sequel to Dubos and Kevlahan (2013), which extends the adaptive wavelet approach to the sphere and reimplements the algorithm in Fortran 95 and message passing interface (MPI) with the goal of achieving high computational efficiency and good parallel scaling.

After we introduce the general numerical model and algorithm in sections 2 and 3, section 4 deals with the technical challenges and modifications of the algorithm due to the non-uniform discrete geometry on the sphere (e.g. the fact that triangular cells are no longer uniform or equilateral). The parallel implementation, data structure and strategies for optimizing computational efficiency are described in section 5. This section also summarizes the strong and weak parallel scaling performance of the method. Sections 6 and 7 verify the accuracy of the model for standard smooth test cases (Williamson *et al.*, 1992) and for a more complex nonlinear test case (Galewsky *et al.*, 2004). Finally, we consider the most challenging shallow-water test case for dynamically adaptive methods: fully developed homogeneous rotating turbulence on the sphere.

#### 2. Wavelets on the sphere

# 2.1. Wavelet spaces

A function f(x) defined on a domain  $\Omega \subset \mathbb{R}^n$  may be approximated by a set of discrete basis functions  $\phi_k^j(x)$ ,

$$f(x) \approx \sum_{k \in \mathcal{K}(j)} f_k^j \phi_k^j(x),$$

where *j* is the scale, *k* is the position,  $\mathcal{K}(j)$  is the index set of positions defining the basis functions at each scale *j* and  $f_k^j$  are the weights (called scaling coefficients). The larger the scale *j*, the finer and more accurate the approximation and the bigger the index set  $\mathcal{K}(j)$ . Alternatively, we can represent f(x) in wavelet space in terms of the difference between successive levels of approximation *j* and j + 1, which is spanned by the set of wavelet functions  $\psi_m^j(x)$ ,

$$f_{J_{\max}}(x) = \sum_{k \in \mathcal{K}(J_{\min})} f_k^J_{\min} \phi_k^J_{\min}(x) + \sum_{j=J_{\min}}^{J_{\max}-1} \sum_{m \in \mathcal{M}(j)} \tilde{f}_m^j \psi_m^j(x),$$
(1)

where  $\mathcal{M}(j)$  is the index set of positions defining the wavelets at each scale *j*. Note that we require a coarse representation at scale  $J_{\min}$  and we have truncated the representation at a finest level of resolution  $J_{\max}$ . The basis functions  $\phi_k^j(x)$  spanning each scale *j* are called *scaling functions* and the functions  $\psi_m^j(x)$  spanning the difference space between representations at successive scales *j* and *j* + 1 (i.e. the interpolation error) are called *wavelets*.

This wavelet multiresolution analysis relies on the fact that the grids at two successive scales j and j + 1 are *nested*. The index sets  $\mathcal{K}(j)$  and  $\mathcal{M}(j)$  refer to nodes on the grid, and hence each wavelet and scaling function  $\psi_m^j(x)$  or  $\phi_k^j(x)$  (and accordingly each coefficient  $\tilde{f}_m^j$  or  $f_k^j$ ) is uniquely associated with a particular node. Due to the nesting property of the grids, the union of the index sets  $\mathcal{K}(j)$  and  $\mathcal{M}(j)$  equals the index set of nodes at the finer scale j + 1,

$$\mathcal{K}(j+1) = \mathcal{K}(j) + \mathcal{M}(j).$$

This relation reflects the fact that the wavelets span the difference in approximation spaces between successive scales j and j + 1. It is also the basis of adaptive wavelet methods, since the wavelet coefficients  $\tilde{f}_k^j$  measure directly the interpolation error associated with deleting a node  $x_k^j$  from the grid. Wavelet-based adaptivity is described in detail in section 2.3.

Figure 1 shows three levels of nested grids j = 0, 1, 2 on the sphere. The round (blue in the online article) nodes are a coarse



**Figure 1.** Nested grids on the sphere. An icosahedron projected on to the sphere forms a coarse approximation (i.e. the round (blue in the online article) nodes denoted by index set  $\mathcal{K}(0)$ ).Refining this coarse grid via edge bisection and projection on to the sphere produces a finer nested grid.  $\mathcal{M}(0)$  are the new nodes added between level j and j + 1 and the finer grid has index set  $\mathcal{K}(1) = \mathcal{K}(0) + \mathcal{M}(0)$ . Note that wavelets are located at nodes given by the index sets  $\mathcal{M}(j)$  and scaling functions are located at nodes with index sets  $\mathcal{K}(j)$ . The triangles (red in the online article) are the new points added at the next finer level of approximation j = 2.

grid (level 0, index set  $\mathcal{K}(0)$ ). Together with the square (green in the online article) nodes  $\mathcal{M}(0)$ , they give the next finer level j = 1and satisfy the nested property  $\mathcal{K}(0) + \mathcal{M}(0) = \mathcal{K}(1)$ . Similarly, by adding the triangles (red in the online article) we construct the next finer approximation level j = 2 consisting of all nodes of any colour or shape. Accordingly, in level j,  $\mathcal{K}(j)$  are all the node positions obtained from refining the icosahedron j times and  $\mathcal{M}(j)$  are all the edge bisections on level j.

# 2.2. Discrete wavelet transform

Second-generation wavelets (Sweldens, 1996) allow the computation of the wavelet coefficients  $\tilde{f}_m^j$  from the scaling coefficients  $f_k^{j+1}$  by a discrete wavelet transform referred to as *lifting*. Starting on the finest level  $J_{\text{max}}$  and working successively down to the coarsest level  $J_{\text{min}}$ , one computes

$$\tilde{f}_m^j = f_m^{j+1} - \sum_{k \in \mathcal{K}_m^j} \tilde{s}_{km}^j f_k^{j+1} \quad \forall m \in \mathcal{M}(j).$$
(2)

This is called the *predict* step, since the last term predicts (using interpolation) the values of the scaling coefficients  $f_m^{j+1}$  that will be neglected at the coarser scale *j*. The  $\tilde{f}_m^j$  are the wavelet coefficients and they measure the local interpolation error at scale *j*. Finally, one *updates* the scaling function coefficients at scale *j* by adding a linear combination of the neighbouring wavelet coefficients:

$$f_k^j = f_k^{j+1} + \sum_{m \in \mathcal{M}_k^j} s_{km}^j \tilde{f}_m^j \quad \forall k \in \mathcal{K}(j).$$
(3)

This update step is used to improve the properties of the transformation. In our case, we design the update step to ensure that the mean is conserved during refinement (i.e. prolongation) or coarsening (i.e. restriction) between different levels of resolution. Note that the predict and update weights  $\tilde{s}_{km}^{j}$  and  $s_{km}^{j}$  are zero except in a small neighbourhood of *l* or *k*, respectively, i.e. they have finite and compact stencils.

Since, unlike first-generation wavelets, second-generation wavelets are constructed in physical space, they can be designed for irregular domains and curved geometries. Second-generation wavelets were first developed for the sphere by Schröder and Sweldens (1995). The nested grid is generated by repeatedly bisecting the edges of an icosahedron, which forms the level j = 0.

(The dots in Figure 1 are the vertices of an icosahedron.) Each bisection increases the scale *j* by one. In practice, the coarsest level for the wavelet transform  $J_{\min} > 0$ , since the icosahedral grid is far too coarse to be useful in practice. As explained in section 3.2, the position of the grid points obtained must be adjusted slightly, since after many such bisections the resulting triangular cells are increasingly non-uniform and far from the ideal case of equilateral triangles (at least near the 12 vertices and 30 edges of the original icosahedron). We employ grid-improvement techniques that optimize globally the geometrical properties that are important for the accuracy of the Thuburn-Ringler-Skamarock-Klemp (TRiSK) finite-volume/finite-difference scheme we use to approximate the shallow-water equations.

We now describe how filtering the wavelet coefficients can make this nested multiscale grid dynamically adaptive.

# 2.3. Adaptivity using wavelets

The discrete approximation of the function f(x) at the finest scale  $f_{j_{max}}$  can be compressed by removing (i.e. setting to zero) all those wavelet coefficients with modulus below a specified tolerance threshold  $\varepsilon$ . Due to the one-to-one correspondence between wavelet coefficients  $\tilde{f}_m^{j}$  and grid points  $x_m^{j}$ , an adapted grid is obtained by including all grid points that correspond to active wavelet coefficients. Since the wavelet coefficients are exactly the local interpolation error, this filtering ensures that the error of the compressed function constructed by inverse wavelet transform on the adapted grid is at most  $\varepsilon$ .

In addition to those significant wavelet coefficients above the threshold,  $|\tilde{f}_m^j| \ge \varepsilon$ , the adapted grid also includes all grid points on the coarsest level  $J_{\min}$  and all grid points that are adjacent in space (i.e. on the same level j) or in scale (on the next finer level j + 1) to the significant wavelet coefficients. This allows dynamic adaptivity, since adding nearest neighbours allows the grid to track energetic features as they move or develop smaller scales over one time step. This basic approach to wavelet adaptivity was first proposed by Liandrat and Tchamitchian (1990). For more details on wavelet-based adaptive numerical methods for partial differential equations, we refer the reader to the review by Schneider and Vasilyev (2010).

# 3. Conservative wavelet method for the shallow-water equations on the sphere

# 3.1. Discrete shallow-water equations and flux restriction

The evaluation of the free surface height perturbation  $\delta h$  and horizontally averaged velocity **u** of a thin layer of fluid is described by the vector-invariant rotating shallow-water equations:

$$\frac{\partial \delta h}{\partial t} + \operatorname{div} F = 0, \tag{4}$$

$$\frac{\partial \mathbf{u}}{\partial t} + qF^{\perp} + \operatorname{grad} B = 0, \tag{5}$$

with potential vorticity

$$q = \frac{\operatorname{curl} \mathbf{u} + f}{h},$$

thickness flux  $F = h\mathbf{u}$ , height  $h = H + b + \delta h$ , Bernoulli function  $B = g \,\delta h + K$ , kinetic energy  $K = |\mathbf{u}|^2/2$ , Coriolis parameter f, bottom topography b, mean height H and gravitational acceleration g.  $F^{\perp}$  is the flux perpendicular to the thickness flux F.

All differential operators are discretized using second-order finite volumes or finite differences as described in Ringler *et al.* (2010) and the energy-conserving variant is chosen for  $qF^{\perp}$ . The prognostic variables  $\delta h$  and u are arranged in a staggered fashion: the scalar values h are located on nodes of the triangular grid and

the vector field **u** is discretized by storing the normal velocity at the edge midpoints of the triangular cells located at the edge bisection. There are therefore three velocity components associated with each height variable and they are oriented parallel to each edge in a counter-clockwise fashion. Thus, each triangular cell is associated with four discrete prognostic variables. Since we have two sets of variables on two different grids, we require two distinct wavelet transforms: a scalar transform for height h and a vectorvalued wavelet transform for the velocity u. These transforms are described in detail in Dubos and Kevlahan (2013). For the implementation, the prognostic variables  $\delta h$  and **u** can be stored as either scaling coefficients (i.e. in physical space) or wavelet coefficients (i.e. in wavelet space), since the wavelet transforms allow us to convert between the two at any time. While Dubos and Kevlahan (2013) chooses to represent in wavelet space, the implementation of this model stores scaling coefficients, which leads to rearrangements in the algorithm and saves some operations because values are accessed more frequently in physical space.

The computation of one time step begins with computing the intermediate values B, F and  $qF^{\perp}$  by the TRiSK operators on the locally finest level and restricting them to coarser levels until they are available on all active elements. The computations are preceded by setting masks for different quantities. This includes the previous height values, which are computed by the inverse wavelet transforms as needed. This way, every operator has all values available that are required for its evaluation. Finally, the gradient and divergence operators are evaluated and new  $\delta h$  and u variables are computed from the trends. At this point, the time step is not yet completed. Partial wavelet transforms (for height only) need to be computed to ensure consistency and mass conservation between levels, since prognostic variable are stored as scaling coefficients. The grid is then adapted based on the wavelet coefficients. This means keeping, or if necessary adding, all grid points that either correspond to an active coefficients or are needed for the stencil of one of the differential operators, and removing the rest. Since there are two types of wavelet coefficient (for h and  $\mathbf{u}$  respectively), the adaptation step also includes a consistency step that guarantees that active *h* grid points (nodes) have active **u** points (edges) in their vicinity and vice versa.

The wavelet coefficient tolerance  $\varepsilon$  defined in section 2 is not the actual threshold used for grid adaption. Instead, it is a parameter that is set in order to control the error in the trend. In turn,  $\varepsilon$  determines the actual tolerances  $\varepsilon_h$  and  $\varepsilon_u$  on the height and velocity wavelet coefficients. The relation between the thresholds for velocity wavelet coefficients,  $\varepsilon_u$ , and height wavelet coefficients,  $\varepsilon_h$ , and the trend tolerance  $\varepsilon$  depends on the regime and details can be found in Dubos and Kevlahan (2013).

• Quasi-geostrophic regime:  $\text{Ro} = U/fL \ll 1$ ,

$$\varepsilon_{\rm h} = \frac{f U L \operatorname{Ro}}{g} \varepsilon^{3/2},$$
$$\varepsilon_{\rm u} = U \operatorname{Ro} \varepsilon^{3/2}.$$

• Inertia–gravity regime:  $T \sim L/c$ ,

$$\varepsilon_{\rm h} = \frac{cU}{g} \varepsilon^{3/2},$$
$$\varepsilon_{\rm u} = U \varepsilon^{3/2}.$$

Here *U* and *L* are typical velocity and horizontal length-scales, *c* is the wave speed and Ro is the Rossby number.

For the discretization in time, we use a four-stage third-order strong stability-preserving Runge–Kutta method that is stable up to Courant–Friedrichs–Lewy (CFL) numbers of 2 (Spiteri and Ruuth, 2002). The time-step size is computed depending on the solution to guarantee stability:

$$\Delta t = \min\left(\frac{1}{\omega_{\max}}, \left(\frac{\Delta x}{|\mathbf{u}|}\right)_{\min}\right),$$

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Table 1. Number of bisection refinement levels *J* of the icosahedron, number of height nodes *N*, total degrees of freedom (d.o.f.), average edge length  $\Delta x$  and equivalent truncation limit for spherical harmonic spectral solvers *T*. Level *J* = 12 corresponds to a resolution of approximately 1 arcmin on the Earth.

J	Ν	d.o.f.	$\Delta x  (\mathrm{km})$	Т
3	642	2562	959.3	13
4	2562	10 242	479.6	25
5	10 242	40 962	239.8	51
6	40 962	163 842	119.9	101
7	163 842	655 362	60.0	202
8	655 362	2621 442	30.0	404
9	2621 442	10 485 762	15.0	809
10	10 485 762	41 943 042	7.5	1619
11	41 943 042	167 772 162	3.7	3238
12	167 772 162	671 088 642	1.9	6476

where  $\omega_{\text{max}} = \left(\sqrt{f^2 + gh\pi/\Delta x}\right)_{\text{max}}$  is the maximum frequency supported on the grid.

The maximum level  $J_{max}$  may be determined implicitly by the tolerance  $\varepsilon$  or it can be set explicitly. Allowing  $J_{\text{max}}$  to be set by  $\varepsilon$  ensures spatially homogeneous error, but since an additional level is always added it also adds computational overhead. It is also sometimes preferable to know the minimum grid resolution in advance of the simulation. Similarly, the choice of the coarsest level  $J_{\min}$  also affects the efficiency of the method, since retaining several completely filled levels results in unnecessary wavelet transform steps. (By a filled level, we mean that the grid adaptation criteria force all grid points on a particular level to be retained.) Thus, efficiency requires that  $J_{\min}$  should not be less than the highest filled level. Furthermore, if a particular level j is almost entirely filled, it is still preferable to set  $J_{\min} \ge j$ , since the extra nodes added are compensated by the gains of removing the lower level(s). Although the choice of  $J_{\min}$  does not affect accuracy directly, increasing the minimum level can improve accuracy indirectly by improving grid quality (see section 3.2). Typical values used in the test cases described below are  $J_{\min} = 6$  (i.e. six dyadic refinements of the icosahedron) for a localized test function and  $J_{\min} = 7$  for a global test function if there are  $\mathcal{O}(10^6)$  degrees of freedom (d.o.f.) in the adaptive model. Table 1 shows different grid sizes and compares them with the equivalent spherical harmonic truncation limit T.

The model described above is inviscid and the only source of dissipation is due to the wavelet adaptivity. We also consider the shallow-water equations with explicit dissipative terms added to both the height (4) and velocity (5) equations:

$$\frac{\partial \delta h}{\partial t} + \operatorname{div} F = \nu \operatorname{div} \operatorname{grad}(\delta h), \frac{\partial \mathbf{u}}{\partial t} + qF^{\perp} = -\operatorname{grad}(B+K) + \nu \left(\operatorname{grad}\operatorname{div}\mathbf{u} - \operatorname{grad}^{\perp}\operatorname{curl}\mathbf{u}\right),$$

where the new parameter  $\nu$  is the viscosity. The viscosity can be chosen to limit the minimum length-scale or to model dissipative mechanisms in the ocean or atmosphere (e.g. subgrid-scale turbulent viscosity).

#### 3.2. Grid optimization

Since the discretization of the differential operators from Ringler *et al.* (2010) is second-order accurate for equilateral triangles but drops to first-order accurate when the triangles are far from equilateral, optimizing grid quality improves the accuracy of the solutions. As explained earlier, this optimization is especially important for large number of scales (e.g. approximately for scales J > 6), since the grid becomes increasingly distorted near the edges of the original icosahedron as the grid is successively refined by edge bisection.

Figure 2 shows a section of the grid obtained by simple edge bisection (top) and the same section for an optimized



**Figure 2.** Grid quality of simple bisection (top) and an optimized grid (bottom). The offset (red in the online article, in metres at J = 7) of edge bisection of primal (black) and dual (green in the online article) grids is reduced when the grid is optimized using the method of Heikes *et al.* (2013). The offset error has been reduced by a factor of about 60. This results in a more accurate discretization of the differential operators in the TRiSK scheme.

grid (bottom) at J = 7 obtained using the method of Heikes *et al.* (2013). The approximation of the Laplacian operator is guaranteed to converge if the bisection of the primal edge (black triangles) and dual edge (hexagons, green in the online article) coincide. The distance (marked in red in the online article) between those two intersection points is an important measure for the grid quality. On simple grids refined by edge bisection, the Laplacian operator does not even achieve first-order convergence.

Optimized grids provided by Heikes *et al.* (2013) can be read into the model (this is currently the default method used in this article). This approach seems to provide the best optimization and leads to a convergent Laplacian operator. As an alternative, the grid optimization proposed by Xu (2006) has also been implemented. Advantages of this method are that it optimizes locally (rather than globally), is computationally inexpensive and is easy to implement. However, while the grid quality is improved, leading to lower error for a given resolution *J*, the Laplacian operator does not converge. In both cases, the grid is first optimized on a coarsest level  $J_{min}$  (determined by the physics of the problem and computational resources). Finer levels  $j > J_{min}$  are obtained by edge bisection. This is necessary because the interscale restriction and prolongation operators used in the adaptive wavelet method require the grid points to be nested.

# 4. New challenges from spherical geometry

# 4.1. General issues

The main contributions of this work are to extend the planar model of Dubos and Kevlahan (2013) to allow for a non-uniform grid of non-equilateral triangles and to develop a highly efficient parallelized code and associated data structure. In this section, we consider the special challenges arising from the non-uniform discrete C-grid on the sphere. In particular, due to fact that the weights and stencil geometry for discrete differential operators depend on position, the hexagonal grids on successive levels have complicated overlap regions and the convergence behaviour of operators is affected. In the present method, all areas are computed as spherical polygons, edges are spherical arcs and lengths are computed as arc lengths on the sphere. In contrast to the plane, where only hexagonal cells occurred in the dual grid to the triangular primal grid, the sphere includes 12 exceptional pentagonal cells, corresponding to the 12 vertices of the original icosahedron.

As in the planar version, the velocity requires a non-separable vector-valued wavelet transformation. This transformation involves interpolating the velocity at the mid-point of an edge of a fine-level triangle j + 1 from values of the edges on the coarser level *j*. Interpolation is carried out as a linear combination of the velocity values on the coarse edges, where the (local) weights are

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pre-computed to guarantee second-order accuracy. At least six edges are required theoretically for second-order accuracy, but the model uses 12 edges in a symmetric stencil to gain stability and higher accuracy (see Dubos and Kevlahan, 2013 for more details). On the sphere, every edge needs to compute and store its own weights, which are obtained by solving two  $6 \times 6$  systems of linear equations.

Combining Eqs (2) and (3) gives the action of the heightrestriction operator  $R_{\rm h} h_k^{j+1} = h_k^j$  as

$$h_{k}^{j} = h_{k}^{j+1} + \sum_{m \in \mathcal{M}_{k}^{j}} s_{km}^{j} h_{m}^{j+1} - \sum_{m \in \mathcal{M}_{k}^{j}} \sum_{k' \in \mathcal{K}_{m}^{j}} s_{km}^{j} \tilde{s}_{k'm}^{j} h_{k'}^{j+1}.$$
 (6)

The scaling function coefficient at node k on level j,  $h_k^j$ , corresponds to the average height on the hexagon with centre node k. The filter coefficients  $\tilde{s}$  and s are chosen such that the restriction from level j + 1 to level j conserves total height (i.e. conserves mass):

$$\sum_{k \in \mathcal{K}(j)} A_k^j h_k^j = \sum_{k \in \mathcal{K}(j+1)} A_k^{j+1} h_k^{j+1} \left( = \int_{\text{Sphere}} h \right), \quad (7)$$

where *h* is the continuous height, defined as the result of an inverse wavelet transform to infinite resolution. Using Eqs (3) and (2) to express  $h_k^{j+1}$  from  $h_k^j$  and  $\tilde{h}_m^j$ , then setting to zero all but one coefficient from among  $h_k^j$  and  $\tilde{h}_m^j$  yields the following conditions:

$$A_m^{j+1} = \sum_{k \in \mathcal{K}_m^j} s_{km}^j A_k^j,$$
  
$$A_k^j = A_k^{j+1} + \sum_{m \in \mathcal{M}_k^j} \tilde{s}_{km}^j A_k^{j+1}.$$

These conditions are satisfied by letting

$$\tilde{s}_{km}^{j} = \frac{A_{km}^{j+1}}{A_{m}^{j+1}}, \qquad s_{km}^{j} = \frac{A_{km}^{j+1}}{A_{k}^{j}}$$

where  $A_{km}^{j+1}$  is the area shared by the coarse-level hexagon  $A_k^j$  and the fine-level hexagon  $A_m^{j+1}$  (see Figure 3). Note that partial areas  $A_{km}^{j+1}$  cover the fine- and coarse-scale hexagons, ensuring that  $s_{km}^j$  and  $\tilde{s}_{km}^j$  are indeed weights. Thus, it is necessary to compute the areas of intersection of spherical polygons. Hexagonal cells (and pentagons) are subdivided into six (or five) triangles using the central point (i.e. the barycentre). Since the types of triangle intersections that can appear during the  $A_{mk}^j$  computation are only a subset of all possible intersection cases, the intersection computation is optimized to account only for cases that can occur. The points of intersection of triangle edges are computed as spherical arc (great circle) intersections.

As in the planar case (Dubos and Kevlahan, 2013), in order to guarantee mass conservation the fluxes need to be restricted and the restriction operators must satisfy the commutation relation:

$$R_{\rm h} \circ \operatorname{div}_{i+1} = \operatorname{div}_i \circ R_{\rm F}.$$
(8)

On the sphere the construction of a flux-restriction operator  $R_F$  that guarantees this commutation property for a given height-restriction operator  $R_h$  poses additional difficulties, due to location-dependent discrete geometry and the problem of overlapping hexagons at successive levels described above.

We use the strategy proposed by Dubos and Kevlahan (2013) to split the height- and flux-restriction operators into a basic and correction part:

$$R_{\rm F} = R_{\rm F0} + R'_{\rm F}.$$
 (9)

The more complicated basic and corrective flux restrictions  $R_{F0}$  and  $R'_f$  needed in spherical geometry are described in the following two subsections.



**Figure 3.** Arrangement of fine- and coarse-scale height nodes used in the calculation of the corrective flux restriction through coarse edge  $kl_{2j}$  (indicated by the arrow). The figure also shows the two partial areas  $A_{km}^{j+1}$  and  $A_{lm}^{j+1}$  used in the calculation of the flux restriction  $(A_m^{j+1})$  is the complete fine-scale hexagon (green in the online version) with partial areas  $A_{km}^{j+1}$  and  $A_{lm}^{j+1}$ ).

### 4.2. Basic flux restriction

In the following, notation from Figure 4 will be used, where all quantities (particularly u, v, w, x and F) are integrated fluxes (total flux through an edge or part of an edge), except that A stands for area. As shown in Figure 4, the total area A of the central (green in the online article) fine-scale hexagon is decomposed as  $A = A_1 + A_2 + A_3 + A_4$  according to the way it overlaps with the two adjacent coarse-scale hexagons (red in the online article) sharing the solid (red) edge.

We assume that we are given all fluxes u, v, w on the fine grid (green) and want to compute the flux through the solid (red) coarse edge  $F = F_1 + F_2$  shown in Figure 4.  $F_2$  is the flux through the part of the coarse edge outside the (green) fine hexagon and  $F_1$  is the flux through the part of the coarse edge inside the (green) fine hexagon. Here we consider the case where one end of the coarse edge is inside the fine (green) hexagon and the other end is outside. In this way, the procedure for both cases (ending inside and ending outside) is explained. In the case in which both or no edges are inside the fine hexagon, the same procedure is simply applied at both ends.

The sum of fluxes entering the fine (green) hexagon on the left of the  $F_1 + x_{-1}$  connection is defined as  $F_{in}$  and the flux leaving on the right is defined as  $F_{out}$ :

$$F_{\rm in} = -w_{-1,1} + u_{-1,0} - v_{-1,-1} + w'_{-1,1}, \qquad (10)$$

$$F_{\text{out}} = -v_{1,1} + u_{1,0} - w_{-1,1} + w'_{-1,1}.$$
 (11)

The divergence theorem says that the average divergence of a vector field over an area A, div<sub>A</sub>, is equal to the net flux through the boundary of A divided by A,

$$\operatorname{div}_{A} = \left(F_{\operatorname{out}} - F_{\operatorname{in}}\right) / A. \tag{12}$$

Therefore, average divergence over the small area shown in Figure 4 may be written as

$$\operatorname{div}_{A_1+A_2} = \frac{(F_1 + x_{-1}) - F_{\text{in}}}{A_1 + A_3} = \frac{F_{\text{out}} - (F_1 + x_{-1})}{A_2 + A_4}$$

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Figure 4. Small overlap regions between hexagons at successive levels need to be accounted for when restricting the thickness flux. In the online article, the coarse-level hexagons are red and the fine-level hexagons are green. The inset figures show close-up views of the small overlapping areas due to the non-uniform C-grid structure on the sphere.

The above expression can be solved for the flux  $F_1 + x_{-1}$ , using  $A_2 + A_4 + A_1 + A_3 = A$ , to find

$$F_1 + x_{-1} = \frac{F_{\text{in}} (A_2 + A_4) + F_{\text{out}} (A_1 + A_3)}{A}$$

An expression similar to Eq. (12) also holds for the small triangle associated with area  $A_{-1}$ ,

$$A_{-1}\operatorname{div}_{A_{-1}} = -F_2 - w'_{-1,1} - x_1.$$

Solving for  $F_2$  yields an expression for the flux  $F_2$ ,

$$F_2 = -A_{-1} \operatorname{div}_{A_{-1}} - w'_{-1,1} - x_1.$$

Combining these results,  $w'_{-1,1}$  cancels and we find that the total basic restricted flux  $F_0$  corresponding to the action of the operator where  $R_{\rm F0}$  on the fine-scale fluxes is

$$F_{0} = F_{1} + F_{2}$$

$$= \frac{A_{2} + A_{4}}{A} \left( -w_{-1,1} + u_{-1,0} - v_{-1,-1} \right)$$

$$+ \frac{A_{1} + A_{3}}{A} \left( -v_{1,1} + u_{1,0} - w_{-1,1} \right)$$

$$- A_{-1} \operatorname{div}_{A_{-1}} - x_{1} - x_{-1}.$$
(13)

The remaining step is to compute the fluxes  $x_1$  and  $x_{-1}$ through the small boundaries shown in the zooms in Figure 4. The flux through boundary  $x_1$  is interpolated using the fluxes at fine edges on the upper half  $u_{0,1}$ ,  $v_{1,1}$ ,  $w_{-1,1}$ ,  $u_{-1,0} - v_{-1,1}$ ,  $v_{-1,2} - v_{-1,1}$  $w_{1,2}, w_{1,1} - u_{1,0}$ . The flux on the lower half through  $x_{-1}$  is found in the same way from the fluxes  $u_{0,-1}, w_{-1,1}, v_{-1,-1}, u_{-1,0}$  $w_{-1,-1}, w_{-1,-2} - v_{1,-2}, v_{1,-1} - u_{1,0}$ . We employ the interpolation formula used for interpolating velocities in Dubos and Kevlahan (2013), which has the following advantages:

- (1) second-order accurate;
- (2) reliable in the case of equilateral triangles; and
- (3) computationally efficient, as it reuses components.

Note that the commutation relation (Eq. (8)) is satisfied irrespective of the interpolation formula used to compute the fluxes  $x_1$  and  $x_{-1}$ .

This completes the computation of the restricted flux obtained from the basic operator  $R_{\rm F0}$  in Eq. (9). We now explain how to compute the corrective part  $R'_{\rm F}$  of the flux restriction in Eq. (9), in order to obtain the full flux-restriction operator  $R_{\rm F}$ .

#### 4.3. Corrective flux restriction

The operator  $R'_{\rm F}$  that guarantees the commutation property (Eq. (8)) can be computed from the hexagon intersection areas above, where F' is the part of the restricted flux obtained from the corrective operator  $R'_{\rm F}$ .

We assume that the hexagonal cell *k* has *N* edges (where N = 5for pentagons and N = 6 for hexagons). The nearest-neighbour fine-scale neighbours are denoted by  $m_0, m_2, \ldots, m_{2N-2}$  and the second nearest neighbour fine-scale neighbours are denoted by  $m_1, m_3, \ldots, m_{2N-1}$ . The nearest-neighbour coarse-scale neighbours are denoted by  $l_0, l_2, \ldots l_{2N-2}$  and the second nearest neighbour coarse-scale neighbours are denoted by  $l_1, l_3, \ldots, l_{2N-1}$ . They are arranged in such a way that the following holds:

- $m_{2j}$  is the midpoint of the edge joining nodes k and  $l_{2j}$ ; the second nearest coarse neighbours of  $m_{2i}$  are  $l_{2i-2}$ ,  $l_{2i+2}$ ;
- $m_{2j+1}$  is the midpoint of the edge joining nodes  $l_{2j}$  and  $l_{2j+2}$ ; the second nearest coarse neighbours of  $m_{2j+1}$  are k and  $l_{2i+1}$ .

The arrangement of the nodes, points and edges used the calculation of the corrective flux restriction is shown in Figure 3.

Using the notation in Figure 3, the definition of the height restriction R<sub>h</sub> and the relation between cell areas at coarse and fine scales, the corrective part of the restricted flux F' for cell k is given by

$$F' = \sum_{j=0}^{N-1} (km_{2j}l_{2j}) + (km_{2j+2}l_{2j}) + (km_{2j-2}l_{2j}) + (km_{2j+1}l_{2j}) + (km_{2j-1}l_{2j}) + \frac{1}{2}(km_{2j+1}l_{2j+1}) + \frac{1}{2}(km_{2j-1}l_{2j-1}) + \frac{1}{2}(l_{2j+4}m_{2j+2}l_{2j}) + \frac{1}{2}(l_{2j-4}m_{2j-2}l_{2j}),$$
(14)

$$(kml) = \frac{A_{km}^{j+1}A_{lm}^{j+1}}{A_m^{j+1}} \left(c_k^{j+1} - c_l^{j+1}\right)$$
(15)

and  $c_k^{j+1} = \operatorname{div}_k^{j+1} F_k^{j+1}$  is the divergence of flux on the fine grid.

In summary, the restricted flux  $R_F F_k^{j+1} = F_k^j$  is found by adding the basic restricted flux found using Eq. (13) to the corrective flux restriction found using Eqs (14) and (15). Note that to find the corrective flux restriction we must first calculate the local areas  $A_k^j$  and  $A_{km}^{j+1}$  associated with all active height nodes  $x_k^{j+1}$  on the fine grid. Using the height restriction (Eq. (6)), it is relatively straightforward to verify that the complete flux restriction defined in Eqs (13) and (14) satisfies the commutation relation (Eq. (8)).

#### 5. Implementation and performance

#### 5.1. General considerations

The algorithm, which was previously implemented in Matlab for planar geometry by Dubos and Kevlahan (2013), has been completely reimplemented in Fortran 95 with the goal of producing a code that is computationally efficient and scales well for parallel computation on large numbers of central processing unit (cpu) cores. We have also made changes to the algorithm itself: the prognostic variables are stored in physical space instead of in wavelet space. Since most operators act in physical space, this saves operations.

Because of the irregular geometry, most quantities (lengths, areas, weights, etc.) must be calculated individually for each computational element. Pre-computing these quantities increases memory use (which indirectly increases cpu time), while computing them as needed considerably increases cpu time. We therefore decided which quantities to compute when a node becomes active and which to compute as needed in order to optimize total efficiency. Additionally, quantities for which precision affects the mimetic properties (like mass conservation) are stored in double precision, while values that are already lower accuracy due to truncation error and do not affect the mimetic properties are stored in single precision.

# 5.2. Hybrid data structure

In terms of grid and data structures, the major difficulties arise from the spherical geometry (i.e. non-regular domain) and the locally and dynamically adapted grid. In addition, the grid is staggered, rather than collocated. Data can be associated with either triangles/circumcentres, edges or hexagons/nodes. The goal of this section is to construct a data structure that accommodates these properties and allows efficient computation of the most time-critical parts of the method: the differential and interscale operators.

A naive approach to deal with the triangular staggered grid on a non-regular domain would be to use a data structure where different grid entities are connected via coordinate references. This has the disadvantage that finding second neighbours becomes expensive, additional data (for the references) have to be stored and communicated as the grid changes and it is difficult to keep data locality under control. A better solution is to use a hybrid data structure.

Ignoring the spherical geometry for the moment, the triangular staggered grid can be represented within a regular data structure by grouping one node, two triangles and three edges into one computational element. Then, unfolding the icosahedron, its grid is made up of 20 triangles that can be grouped into ten lozenges see Figure 5, disregarding the refined regions). Therefore, a grid resulting from refining an icosahedron can be divided into ten subgrids, each of which can be stored and accessed in a regular fashion. Note that, at the edges of the lozenges, the two adjacent regular grids of the original icosahedron are rotated with respect to each other. This is dealt with by surrounding the ten lozenge subdomains by ghost/halo cells, where values are not computed but rather copied from their actual locations. Alternatively, the nested levels of the adapted grid could be stored in a quad tree data structure, but computational overhead during the neighbour search would be higher. Neighbours could also be linked via references, increasing the overhead in terms of memory and occasional cleaning and reference updating.

The best way to proceed in our case is to use a hybrid data structure: a combination of regular and irregular grids. The adapted data structure, the irregular part, uses patches as the smallest elements. A patch constitutes a small regular grid. Inside a patch,



**Figure 5.** Hybrid data structure on an icosahedral grid. It is an irregular tree-like data structure with patches (red in the online version) as the smallest element and a regular grid inside each patch. The figure illustrates an example where a small-scale structure in the centre caused adaptive grid refinement.

computations are efficient. A similar hybrid approach was used by Behrens (2009) and Hejazialhosseini *et al.* (2010). In this way, the references can be used to link patches to neighbours in space and scale, without introducing too much additional computational or memory overhead. Since the granularity introduced by the patches involves computational overhead, a patch size that minimizes the total overhead needs to be found. Minimum patch sizes of  $4 \times 4$  or  $8 \times 8$  seem preferable, depending on the structure of the solution. Figure 5 shows an adapted grid with patch size  $4 \times 4$  and  $J_{min} = 0$ .

Figure 6 shows a section of the grid with two  $4 \times 4$  patches (red in the online article) which are located at the edge of a subdomain and therefore have two rows of ghost cells to their left. Elements in a patch are numbered consecutively, so that indices of neighbours on the same patch are offset by plus or minus one if they are in the same row and plus or minus the number of elements of one row (here 4) if they are above or below the element. Since patches are linked with each other, the upper patch in Figure 6 (elements 52-67) and the lower patch (100-115) know each other's offset, i.e. the index of the other patch's first element. In the same way, a patch knows adjacent ghost elements, which are also organized as patches. On the other sides, there could be further patches (not shown) or there might not be any more patches on this level. As an example, consider element 52. The offset of its southern neighbour is computed as 80 (by finding the index 112 through the neighbouring patch and subtracting 52). The same offset can be used for finding the southern neighbour of elements 53-55. Note that the choice of data structure does not affect the computed solution, only computational and memory efficiency. This means that, on a patch, only active elements (as determined by the adaptive wavelet algorithm) are updated.

# 5.3. Serial performance

In this section, we compare the performance of the serial version of our adaptive wavelet method with a similar non-adaptive single-scale implementation of the TRiSK method (Dupos *et al.*, 2014; personal communication) and a standard spectral implementation for the shallow-water equations (Rivier *et al.*, 2002). All calculations were done on the same machine.

Using the non-adaptive TRiSK implementation, a single time step takes  $3.2 \times 10^{-7}$  s per degree of freedom. The TRiSK simulation uses a uniform resolution, corresponding to  $J_{\text{max}} = 8$  levels and 655 362 height nodes ( $2.6 \times 10^6$  total degrees of freedom) in Table 1.

We now compare the performance of the non-adaptive TRiSK code with a similar adaptive wavelet code. The adaptive code has a maximum scale  $J_{\text{max}} = 10$  and uses five levels of refinement,



**Figure 6.** Section of the computational grid with ghost cells on the left. A patch (red in the online version) is a regular grid of elements (green in the online version). Each regular element is made up of of one node, two triangles and three edges.

from J = 6 to J = 10. (The J = 5 grid is first optimized using the method of Heikes et al. (2013), before being used as the coarse level for the wavelet method.) The total number of active height nodes in the  $J_{\text{max}} = 10$  adaptive wavelet method is 500 962, roughly equivalent to the non-adaptive method. This means that the grid compression ratio, which is defined as the proportion of active grid points, is about 21 for the adaptive wavelet method. The adaptive wavelet method is 3.4 times slower per active node than the non-adaptive method. Nevertheless, since the compression ratio is 21, the adaptive wavelet method is still about six times faster than the non-adaptive method in this case. Since the discretizations are identical, this result gives a good estimate of the total computational overhead, due to the multiscale wavelet adaptivity. Note that the overhead due to the wavelet adaptivity increases with the number of levels of refinement. j = 5 refinement levels corresponds to a local refinement of 32 times, which is usually sufficient.

Spectral solvers are considered to be the most efficient nonadaptive solvers (at least for serial implementations) and so give a good lower bound on computational cost. A time step with the spherical harmonics spectral solver SWBOB (Rivier *et al.*, 2002) takes  $2.2 \times 10^{-7}$  s per degree of freedom for a truncation limit *T*341 with 465 124 height nodes. Therefore, we can conclude that the serial adaptive wavelet TRiSK solver is about five times slower per active node than an equivalent spectral solver with a similar number of active height nodes. However, when compression is taken into account, the adaptive wavelet method is about four times faster than the spectral method, but with a maximum resolution about four times finer.

It is important to note that the cpu time per grid point is largely independent of the compression ratio. This is confirmed in Figure 7, which shows that, while the compression ratio (b) varies by a factor of more than three, the cpu time (a) is approximately constant on average. Thus, the computational overhead of the adaptivity should not depend sensitively on the degree of compression.

In conclusion, we find that if the compression ratio achieved is larger than about three times, then the adaptive model will be faster than an equivalent non-adaptive version. As will be seen below (e.g. Figure 7(b)), even for statistically homogeneous flows like turbulence the typical compression ratios achieved are greater than 10-50. Thus, in the serial case we expect that, in addition to achieving a uniform error and finer local resolution, the adaptive wavelet method should be 3-15 times faster than similar non-adaptive methods. In special cases that are naturally very sparse, such as tsunami propagation, the adaptive code could be several hundred times faster than the non-adaptive code.

Parallelization is vital for high performance in large problems and the following two sections explain the parallel algorithm and evaluate its strong and weak parallel scaling performance.

# 5.4. Grid distribution and parallel algorithm

Our goal is to run on at least several hundred cpu cores in parallel with a weak parallel scaling efficiency (see below) of at least 70–80%, in order to assess the potential of our code to run efficiently on an even larger numbers of cores,  $\mathcal{O}(10^3) - \mathcal{O}(10^4)$ , in the future. In particular, we need to identify where the parallel performance bottlenecks are.

Starting from the ten lozenge subdomains shown in Figure 5,  $10 \times 4^{j}$  subdomains can be obtained by dyadic refinement *j* times (i.e. using *j* levels of adaptive resolution). The subdomains are distributed in parallel over several cpu cores, where each core can have several domains. Having several small domains, rather than one large domain, per core can improve cache efficiency through blocking.

In an adaptive simulation, each subdomain will typically have a different number of active nodes and thus require a different amount of communication. The METIS (Karypis and Kumar, 1995) graph partitioner is used to improve load balancing amongst the cores. METIS allows us to assign weights to the graph nodes (representing the subdomains) and graph edges (representing the number of connections between two neighbouring domains). When the load distribution becomes uneven, due to the dynamic adaptivity, the loads can be redistributed during checkpointing.

Every subdomain is extended to hold as many ghost/halo cells as necessary for the various required operators. The values at the ghost cells are communicated as needed. Intracore communication is performed by copying and intercore communication is performed using MPI. During grid adaption, new patches are added as required and grid connectivity between domains is updated (via MPI as necessary). Communications occur at each trend computation and at each grid adaptation step, the latter being less frequent. There is some leeway in the design of the communication pattern, which we use in order to perform as much communication as possible at each grid adaptation step, so that the frequent communications are as light and fast as possible. In addition, critical communications are carried out locally pointto-point rather than using global communication, where possible. Where applicable, communication is non-blocking so that the computations can continue while communication is taking place in the background.

# 5.5. Parallel performance

We quantify parallel performance with respect to both weak and strong scaling efficiency. All calculations are performed on the SHARCNET cluster REQUIN, which has 1541 AMD Opteron cores and a Quadrics Elan4 interconnect. Each processor has two cores and 8 GB of local memory.



Figure 7. Turbulence test case with tolerance  $\epsilon = 5 \times 10^{-2}$ . (a) Cpu time per active grid point and (b) compression ratio based on the maximum scale  $J_{\text{max}} = 10$ .



**Figure 8.** Strong parallel efficiency scaling. Perfectly balanced (solid) and realistic turbulence test case (dashed). N is the total number of degrees of freedom (four times the number of height nodes) and the numbers on the graph are active degrees of freedom per core for each case.

Strong scaling efficiency  $E_S$  is defined as

$$E_{\rm S}=\frac{t_1}{Nt_N}\leq 1,$$

where  $t_1$  is the time to perform a given computation on one core, N the number of cores used and  $t_N$  the time to perform the computation on N cores.  $E_S$  measures how cpu time decreases as the number of cores increases for a *fixed* problem size. Ideally,  $t_N$  should decrease proportionally with increasing N, since the processes can divide the (constant) work. However, in practice, when the portion of the total computation allocated to each core reaches a lower bound,  $t_N$  no longer decreases due to the non-parallelized part of the code (Amdahl's law) or because of communications overhead.

Figure 8 shows the strong parallel scaling efficiency for a perfectly balanced load (solid line) and the turbulence test case (dashed line). As expected, the unbalanced test case has a lower efficiency. The reason for the lower efficiency is explained below. This result suggests that, for both balanced and unbalanced problems, strong parallel efficiency is acceptable for at least 10<sup>2</sup> cpus.

In practice, weak scaling efficiency is a more useful measure, since high-performance codes are intended for large problems. To measure weak scaling efficiency, the computation per core is kept approximately constant as the number of processors is increased. Weak scaling efficiency  $E_W$  is defined as

$$E_{\rm W} = \frac{t_1}{t_N} \le 1$$

where  $t_N$  is the time needed when running on N processors. However, unlike strong parallel efficiency, an efficient parallel code should maintain  $E_W \approx 1$  independently of the number of cores used. Weak scaling is shown in Figure 9 for a balanced test case. It demonstrates that good weak parallel efficiency can be achieved for at least 640 cores. In particular, if there are at least 20 000 height nodes per core, the weak scaling efficiency is 90% on 640 cores. For 1344 height nodes per core, an acceptable efficiency of 70% can still be reached. Scaling for larger numbers of cpus could not be tested, given resource limitations, although these results suggest that the code should have acceptable parallel performance for at least 1000 cores.

The adaptive algorithm requires a large number of communications, although only the interscale (interpolation and restriction) operators require communication with distant cores. Most operators use the results of a previous operator available



**Figure 9.** Weak parallel efficiency scaling. Good performance is demonstrated for up to 640 cores (the maximum tested) with this perfectly balanced test case. Note that, even with only 1344 height nodes per core, this multilevel adaptive method achieves almost 70% parallel efficiency on 640 cores.

on neighbouring nodes. For the TRiSK operators, it is possible to communicate only the prognostic variables if we compute some intermediate quantities on ghost cells. The communications bottlenecks are the interscale operators: flux restriction, velocity interpolation and height interpolation. After fluxes have been restricted from level i + 1 to j, fluxes on level j need to be communicated before restriction from j to j - 1 is possible (and similarly for the interpolation). This not only means that the number of communications grows with the number of levels, it also poses also a more difficult load-balancing problem. Now, in order to avoid processors waiting at the communication step for others to finish, the amount of work on level *j* should be equally distributed amongst the cores for each level *j*. So not only is it desirable to have the same number of total elements on each core, but the elements should ideally be equally distributed at each individual level. This is a significantly more difficult goal to achieve, especially since the multiscale grid structure changes due to grid adaptation after each time step.

This communications bottleneck currently limits efficient *strong* parallel scaling to about  $10^2$  cpus. There is, however, potential for improvement if multiconstraint load balancing is used and/or the parallelization is extended to a hybrid shared/distributed memory approach.

# 6. Verification

In this section we verify the numerical accuracy, convergence and error control of the adaptive wavelet method against several test cases.

We ran test cases 1, 2 and 6 from the standard shallow-water test suite by Williamson *et al.* (1992) for different thresholds and consequentially different numbers of active grid points, in order to investigate convergence. We also show results from the strongly nonlinear barotropic instability test case by Galewsky *et al.* (2004). All test cases use the following physical parameters appropriate for the Earth: gravitational acceleration  $g = 9.80616 \text{ m s}^{-2}$ , radius  $R = 6.37122 \times 10^6 \text{ m}$  and rotation rate  $\Omega = 7.292 \times 10^{-5} \text{ s}^{-1}$ . Longitude  $\lambda$  and latitude  $\theta$  coordinates are related to Cartesian coordinates (x, y, z) by

$$\theta = \arcsin(z/R), \qquad \lambda = \operatorname{atan2}(y/x).$$

6.1. Advection: cosine bell (Williamson test case 1) and smooth bell

This first test case considers only linear advection of a height field by a prescribed velocity. This case is a good test of the grid



**Figure 10.** Errors with respect to analytic solution after 12 days (one rotation around the Earth), for (a) cosine bell and (b) smooth bell.

adaptation routines and grid stability, which is the ability of the adaptation routine to refine only based on the solution and not as a result of previous refinements leading to runaway refinement. The time-independent advecting velocity field is

$$u(\theta, \lambda) = U(\cos\theta\cos\alpha + \sin\theta\cos\lambda\sin\alpha),$$
  
$$v(\theta, \lambda) = -U(\sin\lambda\sin\alpha),$$

with  $U = 2\pi R/(12 \text{ days})$ . Two different initial conditions for the height perturbation are compared: the cosine bell from test case 1 in Williamson *et al.* (1992),

$$h = \frac{H}{2} \left( 1 + \cos(\pi r/L) \right),$$

and a smooth bell inspired by Galewsky et al. (2004),

$$h = He^{r^2/(r^2 - 2L^2)},$$

with

$$r = R \arccos\left(\cos\theta \cos\lambda\right)$$

and H = 1000 and L = R/3. The second initial condition is included because the the cosine bell is only  $C^1$  continuous at r = L. Because our grid adaptivity routine is based on secondorder interpolation, this non-smoothness at the edge of the cosine bell could potentially affect grid stability. Both initial conditions constitute a localized bell that is advected once around the sphere.

Figure 10(a) shows the convergence results for the cosine bell. The convergence of the error for increasing number of grid points corresponds to the expected second-order accuracy. Recall that the number of active grid points is controlled by the tolerance  $\epsilon$ . The results for the smooth bell shown in Figure 10(b) are essentially the same as for the cosine bell.

The grid after one rotation (12 days) with the cosine bell, for a threshold for the trend of  $\epsilon = 0.02$ , is shown in Figure 11. The minimum level has been set to  $J_{\min} = 4$ . The maximum allowed level was set to  $J_{\max} = 10$ , but only levels up to J = 9 are used. This shows that the actual maximum level is set by the tolerance  $\epsilon$  (i.e. the simulation is fully adaptive in scale). The prescribed velocity in this figure goes from right to left. The grid is refined in the centre, where the cosine bell is located, and leaves a trace of a refined grid that gradually dissipates. The smooth bell in fact shows a similar grid structure; grid instability does not seem to be a problem for the non-smooth cosine bell.

# 6.2. Test case 2: steady-state geostrophic flow

The second test case uses the full shallow-water equations. Height h is defined by

$$gh = gH - \left(R\Omega U + \frac{U^2}{2}\right)\cos\theta$$





**Figure 11.** Grid after one rotation with cosine bell in the centre ( $\epsilon = 0.02$ ,  $J_{min} = 4$ ). The maximum level is determined by the adaption routine.

and velocity as

$$u = U\cos\theta$$
,

with  $U = 2\pi/(12 \text{ days})$  and  $gH = 2.94 \times 10^4 \text{ m}^2 \text{ s}^{-2}$ . The flow is in geostrophic balance, so that the exact solution is equal to the initial condition at all times (steady solution). Figure 12(a) shows that the convergence of the global time integration error is approximately first-order accurate. Figure 12(b) and (c) respectively show that the method is second-order accurate in space and that the accumulated error after 12 days is controlled by the tolerance  $\epsilon$ , as expected.

# 6.3. Williamson test case 6: Rossby-Haurwitz wave

Rossby-Haurwitz waves are a standard test case for the full shallow-water equations. Choosing wave number 4, the initial conditions are the non-divergent velocity field

$$u = R\omega\cos\theta + RK\cos^{4-1}\theta \left(4\sin^2\theta - \cos^2\theta\right)\cos 4\lambda,$$
  
$$v = -RK4\cos^{4-1}\theta \sin\theta \sin 4\lambda,$$

with parameters  $K = \omega = 7.848 \times 10^{-6} \text{ s}^{-1}$  and a height chosen to ensure the flow is in geostrophic balance. This initial field rotates without change around the north–south axis.

Since analytical solutions are not available, solutions from the National Center for Atmospheric Research (NCAR) Spectral Transform Shallow-Water Model STSWM at resolution *T*514 are used as a reference. Figure 13 shows that the convergence of the spatial error of the method is indeed approximately second-order for this full shallow-water test case.

## 6.4. Galewsky disturbed jet

The standard test cases above are supplemented by a strongly nonlinear test case proposed by Galewsky *et al.* (2004): a zonal flow with a height disturbance, which leads to an instability that eventually develops into turbulence. As suggested in Galewsky *et al.* (2004), the simulation is first run without the perturbation, to assure that the numerical scheme is able to maintain balance for at least five days. Figure 14 shows the error in height for the first five days for the non-adaptive TRiSK scheme at resolution  $J_{\text{max}} = 7$ 

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**Figure 12.** Test case 2 after 15 days. Errors for (a) height, (b) dependence of grid size on  $\varepsilon$  and (c) error controlled by  $\varepsilon$ .



**Figure 13.** Test case 6. Error of the adaptive wavelet solution compared with the spectral solver STSWM reference solution for the Rossby–Haurwitz wave test case as a function of the number of active grid points.

compared with results from the adaptive method with threshold  $\epsilon$ , chosen so that the total degrees of freedom are comparable to the non-adaptive simulation (6 × 10<sup>5</sup>). These results show that, for a similar number of degrees of freedom and the same discretization scheme, the adaptive wavelet method maintains a significantly lower error (about three times lower).

We now consider the results for the perturbed jet flow after the instability develops. Results for tolerance  $\varepsilon = 5 \times 10^{-3}$ , coarse scale  $J_{\text{min}} = 7$  and finest scale  $J_{\text{max}} = 9$  are shown in Figure 15. This simulation uses about  $2 \times 10^6$  degrees of freedom, for a compression ratio of 5.25. The contours (solid) of the adaptive wavelet simulation nearly overlap with those of a reference simulation with the non-adaptive TRiSK scheme at the finer uniform resolution  $J_{\text{max}} = 10$ , showing that the adaptive wavelet simulation is quite accurate, even for this highly nonlinear time-dependent test case.

We now consider one of the most challenging applications of a dynamically adaptive method: homogeneous isotropic rotating turbulence on the sphere.

#### 7. Rotating shallow-water turbulence on the sphere

# 7.1. Initial condition structure of solution

As a final challenging test case closer to geophysically relevant applications, we consider initial conditions designed to generate



**Figure 14.** Unperturbed zonal jet test case (Galewsky *et al.*, 2004). For similar numbers of active nodes, the adaptive wavelet method maintains a consistently lower error than the non-adaptive TRiSK scheme.

shallow-water turbulence. The coarsest grid is at level  $J_{\rm min} = 5$ and the finest level is determined by the tolerance  $\epsilon$  (it turns out the finest level required is  $J_{\rm max} = 10$ ). Both inviscid and viscous ( $\nu = 10^4$ ) simulations are run with the same tolerance  $\epsilon = 5 \times 10^{-2}$ , corresponding to about  $2 \times 10^6$  degrees of freedom.

The initial condition is made up of several zonal jets similar to the zonal flow in section 6.4, arranged from north to south as shown in Figure 16. Each zonal jet is perturbed to trigger an instability. After two days, vortices form on each of the jets that then interact to generate the approximately homogeneous and isotropic turbulence shown in Figures 17 and 18.

Figures 17 and 18 show the simulation results after 132 h for the inviscid and viscous runs, respectively. Panel (a) shows the relative vorticity and (b) shows the adapted grid. Each grid level is identified by a distinct colour. The most refined regions corresponding to the darkest colours and are located near the intense vorticity filaments that characterize two-dimensional turbulence.

Figure 7 shows that the compression ratio at t = 132 h is about 15 for the inviscid case (17) and 21 for the viscous case (18). It is important to note that at this time the compression ratio is at its lowest level since the turbulence is most intense (compared with both the initial conditions and dissipated flow at later times). Figure 7 also shows that the cpu time per active point remains roughly constant (a), even though the compression ratio changes

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**Figure 15.** Test case with tolerance  $\varepsilon = 5 \times 10^{-3}$  and  $J_{\text{max}} = 9$ . Height perturbation at (a) 2, (c) 4 and (e) 6 h and relative vorticity at (b) 4, (d) 5 and (f) 6 days. The solution of the  $J_{\text{max}} = 10$  non-adaptive reference simulation is dashed, but the lines are mostly indistinguishable.



Figure 16. Initial conditions for (a) zonal velocity and (b) height for the turbulence test case.

significantly when turbulence first develops and then decays again (b). This shows that there is no appreciable computational overhead associated with the degree of grid compression (sparse or dense grids). Not surprisingly, the compression ratio is lowest when the flow is most turbulent. Nevertheless, the viscous adaptive wavelet code is still about four times faster than the spectral code and six times faster than the non-adaptive TRiSK code at this time for an equivalent maximum resolution. This result confirms that adaptive methods can still be advantageous for statistically homogeneous and isotropic flows, like fully developed two-dimensional turbulence.

# 7.2. Energy and spectrum

The total energy E(t) is defined as

$$E(t) = \frac{1}{2} \int gh \left( gh + |u|^2 \right) \, \mathrm{d}S - \frac{1}{2} c_{\text{wave}}^4 A_{\text{S}},$$

where  $A_S$  is the area of the sphere and the wave speed  $c_{wave}$  is

$$c_{\rm wave} = \sqrt{\frac{g}{A_S} \int h \, \mathrm{d}S}.$$

Due to mass conservation,  $c_{wave}$  is constant. Figure 19(a) shows that the total energy for both viscous and inviscid runs first decreases and then stays at about the same level once the turbulence has developed.

The energy spectrum of the turbulent flows can be estimated by interpolating the adaptive results on a uniform grid and using spherical harmonics

$$f = \sum_{l=0}^{N} \sum_{m=-l}^{l} F_{lm} Y_{lm}.$$

The power spectrum is then defined as

$$S_f(l) = \sum_{m=-l}^{l} |F_{lm}|^2$$

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# Adaptive Wavelet Dynamical Core



**Figure 17.** Inviscid shallow-water turbulence with tolerance  $\varepsilon = 5 \times 10^{-2}$  at time t = 132 h. (a) Relative vorticity and (b) adapted grid.



**Figure 18.** Viscous shallow-water turbulence with tolerance  $\varepsilon = 0.05$  and viscosity  $\nu = 10^4$  at time t = 132 h. (a) Relative vorticity and (b) adapted grid.

Figure 19(b) show the spectrum of the rotational part of the velocity  $\omega_{\nu} = \operatorname{curl}_{\nu} u$ . The energy spectrum has a clearly defined power-law range, with a slope of about -2.2.

# 8. Summary, conclusions and perspectives

This article introduces a dynamically adaptive wavelet model for rotating shallow-water equations on the sphere. This model, based on the TRiSK discretization (Ringler *et al.*, 2010), is an extension to spherical geometry of the method developed for the regular C-grid on the plane by Dubos and Kevlahan (2013). The extension to the sphere is based on subdivisions of the icosahedron needed to overcome several challenges to cope with the irregular local C-grid geometry. In addition to the extension to the sphere, the code has also been parallelized using MPI, using a highly efficient hybrid patch-tree data structure. The METIS (Karypis and Kumar, 1995) graph partitioner is used to improve load balancing amongst the cores. The model has been implemented in Fortran 95, in order to optimize computational efficiency.

The current implementation shows good strong parallel efficiency scaling for real test cases up to  $\mathcal{O}(10^2)$  cores and good weak parallel efficiency scaling for load-balanced scenarios for up to at least  $\mathcal{O}(10^3)$  cores. Acceptable parallel scaling to a larger number of cores should be possible if the parallel implementation



**Figure 19.** Turbulence with tolerance  $\epsilon = 5 \times 10^{-2}$  for the inviscid and viscous runs. (a) Total energy minus the total energy at rest. (b) Energy spectrum for the rotational part of the velocity averaged over the interval t = (132-136 h).

is optimized further, for example by using measurement-based multiconstraint load balancing or a hybrid shared/distributed memory approach. Serial computational performance tests showed that the adaptive wavelet code is about three times slower than a non-adaptive TRiSK code and five times slower than a spectral solver per *active* node. This suggests that the adaptive wavelet code should be faster than non-adaptive codes, provided it achieves a grid compression ratio greater than 5. However, the adaptive wavelet code also guarantees spatially uniform error control, which is not possible using non-adaptive methods.

The convergence, accuracy, error control and efficiency properties of the adaptive wavelet method were confirmed using standard smooth test cases from Williamson *et al.* (1992) and a nonlinear unstable zonal jet test case proposed by Galewsky *et al.* (2004). The method was also used to simulate viscous and inviscid fully developed and decaying homogeneous and isotropic shallow-water turbulence. Even in the challenging case of homogeneous turbulence, the adaptive method was able to achieve high compression ratios of up to 15–50 times, due to the fine-scale vorticity filaments that characterize the flow. In this case, the wavelet method is 3–10 times faster than a spectral code with the same number of degrees of freedom. This suggests that the method should be appropriate for high Reynolds number geophysical flows without obvious large-scale sparsity.

Whether or not a dynamically adaptive method is advantageous depends on the physical problem and the goals of the simulation. If the range of active scales of the physical problem is known in advance and can be fully resolved numerically and the goal is high numerical accuracy, then a high-order non-adaptive method (e.g. pseudo-spectral) is ideal. However, if it is impossible to resolve all active scales fully (or if the active scales are not known a priori), it is preferable to use a low-order dynamically adaptive method that tracks the most energetic structures as they move or change scale. In particular, an adaptive method allows rapidly developing small-scale structures (like cyclones or vorticity filaments) to be captured. In this article, we have shown that our wavelet-based method can control numerical error and achieve high rates of compression, even for statistically homogeneous turbulence. In terms of spectra, an adaptive method resolves the full range of active length-scales, i.e. it resolves the most energetic structures at all scales. In this article, all computational nodes are evolved using the same scale-independent time step. In cases where only a small percentage of nodes are at the smallest scales, a scaledependent time step would improve performance significantly (McCorquodale and Colella, 2011) and is straightforward to implement.

To the best of our knowledge, the models in St-Cyr *et al.* (2008) are the only dynamically adaptive methods for the shallow-water equations on the sphere comparable to the one we present here. They analyze an interpolation-based spectral element shallow-water model on a cubed-sphere grid and a block-structured

finite-volume method in latitude–longitude geometry. It is instructive to compare and contrast our wavelet approach with these methods.

In our case, the differential operators are discretized on an icosahedral grid using the TRiSK approximation proposed by Ringler *et al.* (2010) to conserve important mimetic properties of the shallow-water equations. The grid refinement guarantees a spatially uniform pointwise error estimated using wavelet coefficients, while St-Cyr *et al.* (2008) use an empirical refinement criterion. When applied to the Galewsky *et al.* (2004) unstable zonal jet test problem, our method requires roughly four to five times the number of degrees of freedom in order to obtain a similar quality of solution. This is likely due to the fact that the TRiSK scheme uses only second-order accurate approximations of the differential operators, while St-Cyr *et al.* (2008) use fourth-order accurate approximations (at the cost of more computations per degree of freedom).

St-Cyr *et al.* (2008) measure execution time for the adaptive mesh refinement (AMR) finite-volume code with three dyadic refinement levels on 8, 16 and 24 cores. They find that the AMR code is between 3.9 and 2.2 times slower than the fixed resolution code, similar to our overhead result with five refinement levels. However, their strong parallel scaling appears to be weaker than in our case. The AMR code is only about 67% efficient when increasing the number of cores from 8 to 24. In comparison, the adaptive wavelet code is over 95% efficient for the same range of cores and is 60% efficient when comparing execution time on one core with execution time on 640 cores.

The model presented here has been implemented for bisections of an icosahedron. However, the adaptive method, and in particular the interscale operators for flux restriction and height and velocity wavelet transform, can in principle be applied to hierarchical bisections of any Voronoi diagram.

Work is currently under way to incorporate coastlines and variable bathymetry, with the short-term goal of developing a shallow-water global oceans model. This model will be applied to both tsunami propagation and the development and long-term dynamics of ocean flow, such as wind-driven gyres and western boundary currents. In the medium-term, the model will be extended hydrostatically in the vertical direction, while maintaining adaptivity based on horizontal structure. The long-term goal of this work is to evaluate the potential of dynamically adaptive wavelet-based multiscale methods as dynamical cores for the next generation of climate and weather global circulation models.

# Acknowledgements

NKRK acknowledges the support of an NSERC Discovery grant, a Mobility Grant from the French Embassy in Ottawa and a visiting professorship at École Polytechnique. MA acknowledges the support of a Mobility Grant from École Polytechnique.

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